

2.2 Spacetime

The central idea of relativity is that space and time are unified into **spacetime**.

2.2.1 Structure of spacetime

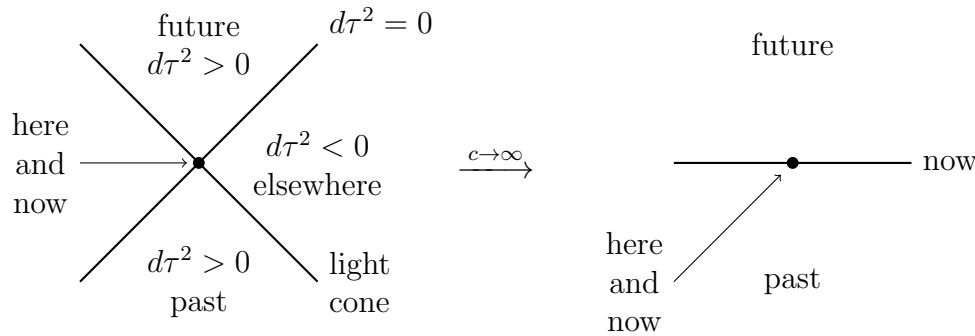


Figure 2.2.1: Relativistic spacetime and its Newtonian limit.

In Minkowski coordinates, an infinitesimal displacement squared can be expressed in terms of the **proper time** τ

$$d\tau^2 = dt^2 - \frac{1}{c^2} (dx^2 + dy^2 + dz^2) \quad (2.2.1)$$

or equivalently in terms of the **proper distance** s

$$ds^2 = dx^2 + dy^2 + dz^2 - c^2 dt^2 \quad (2.2.2)$$

where the minus sign allows us to distinguish time-like and space-like directions, see Figure 2.2.1. Note that only $d\tau^2$ or ds^2 is physical while dt^2 and $dx^2 + dy^2 + dz^2$ are coordinate dependent. In the Newtonian limit these reduce to

$$d\tau^2 \Big|_{c \rightarrow \infty} = dt^2 \quad (2.2.3)$$

$$ds^2 \Big|_{dt=0} = (dx^2 + dy^2 + dz^2)_{dt=0} \quad (2.2.4)$$

2.2.2 Spacetime decomposition

To view things from our usual Newtonian space and time perspective, we introduce a Newtonian time coordinate t which defines a set of spatial hypersurfaces and a corresponding covector field

$$e_{\mathbf{a}}^t = \nabla_{\mathbf{a}} t \quad (2.2.5)$$

We also introduce a spatial rest-frame which defines a fibration of the hypersurfaces by one dimensional time lines and a corresponding vector field $e_{\mathbf{t}}^{\mathbf{a}}$ satisfying

$$e_{\mathbf{t}}^{\mathbf{a}} e_{\mathbf{a}}^t = 1 \quad (2.2.6)$$

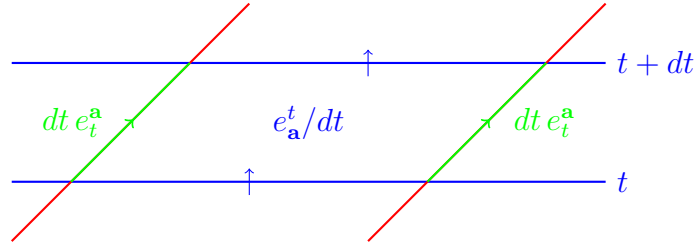


Figure 2.2.2: **Fibration** of **spatial hypersurfaces** and corresponding time vectors e_t^a and covectors e_a^t .

see Figure 2.2.2.

Now we can decompose a spacetime displacement as

$$dx^a = dt e_t^a + dx_3^a \quad (2.2.7)$$

with the spatial displacement satisfying

$$e_a^t dx_3^a = 0 \quad (2.2.8)$$

i.e. spatial displacements lie within the spatial hypersurfaces.

The metric is decomposed as

$$g_{ab} = A e_a^t e_b^t - B_a e_b^t - e_a^t B_b - h_{ab} \quad (2.2.9)$$

with the shift satisfying

$$e_t^a B_a = 0 \quad (2.2.10)$$

and the spatial metric satisfying

$$e_t^a h_{ab} = 0 \quad (2.2.11)$$

giving

$$d\tau^2 = g_{ab} dx^a dx^b = A dt^2 - 2B_a dx_3^a dt - h_{ab} dx_3^a dx_3^b \quad (2.2.12)$$

The inverse metric is decomposed as

$$g^{ab} = \frac{(e_t^a - B^a)(e_t^b - B^b)}{A + B^c B_c} - h^{ab} \quad (2.2.13)$$

where the inverse spatial metric h^{ab} is defined by

$$h^{ab} h_{bc} = \delta_c^a - e_t^a e_c^t \quad (2.2.14)$$

and

$$B^a \equiv h^{ab} B_b \quad (2.2.15)$$

so that

$$e_t^a B^a = 0 \quad (2.2.16)$$

and

$$e_a^t h^{ab} = 0 \quad (2.2.17)$$

In the special case of $B_a = 0$, Eqs. (2.2.9) and (2.2.13) simplify to

$$g_{ab} = A e_a^t e_b^t - h_{ab} \quad (2.2.18)$$

$$g^{ab} = A^{-1} e_t^a e_t^b - h^{ab} \quad (2.2.19)$$